

ASSESSMENT and QUALIFICATIONS ALLIANCE

# Mark scheme June 2003

## GCE

## Physics A

Unit PHA5/W

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#### Units 5 - 9 : Section A

 $\frac{1}{2}$ 

(a)(i)  $^{A}_{Z}X$ C А В correct arrows: A  $\checkmark$ B✓ C✓ (a)(ii)  $e^{-1} + {}^{A}_{Z}X \rightarrow {}^{A}_{Z-1}Y + v_{e} \checkmark$ (4) (b)(i)  $((4.18 - 1.33) \times 10^{-13}) = 2.85 \times 10^{-13}$  (J)  $\checkmark$ (b)(ii)  $1.33 \times 10^{-13}$  (J)  $0.30 \times 10^{-13}$  (J)  $1.63 \times 10^{-13}$  (J) for 3 correct values  $\checkmark$ (b)(iii) (use of  $\Delta E = hf$  gives)  $f\left(=\frac{1.63 \times 10^{-13}}{6.63 \times 10^{-34}}\right) = 2.46 \times 10^{20}$  Hz  $\checkmark$ (allow C.E. from (b)(ii) if largest value taken) (3) (c)(i) ( $\checkmark$  for each precaution <u>with</u> reason to <sub>max</sub>2) handle with (long) (30 cm) tweezers because the radiation intensity decreases with distance store in a lead box (immediately) when not in use to avoid unnecessary exposure to radiation [or any sensible precaution with reason] (b)(ii)  $\gamma$  rays are more penetrating and are therefore more hazardous (to the internal organs of the body)  $\beta^{-}$  particles are more hazardous because they are more ionising  $\checkmark$ 

(  $\checkmark$  for any argued case for either radiation)

 $\frac{(3)}{(10)}$ 

### Unit 5 : Section B

2

(a)	ray diagram to show: rays reflected at concave mirror ✓		
	rays reflected at convex mirror ✓ rays crossing in front of eyepiece ✓	(3)	
(b)	different focal points for rays at different distances from axis $\checkmark$ shortest focal length for paraxial rays $\checkmark$	(2)	
(c)	light of different wavelengths refracted to different foci $\checkmark$ diagram showing refraction with blue focal length closest to lens $\checkmark$	max <u>(2)</u> (7)	

3  
(a)(i) 
$$d = \frac{50 \times 10^6}{3.26} = 15.3 \times 10^6$$
 (pc)  $\checkmark$ 

(a)(ii) (use of 
$$v = Hd$$
 gives)  $v = 65 \times 10^{-6} (\text{km s}^{-1} \text{ pc}^{-1}) \times 15.3 \times 10^{6} \checkmark \approx (1000 \text{ km s}^{-1})$ 

(a)(iii) (use of 
$$\frac{\Delta\lambda}{\lambda} = -\frac{v}{c}$$
 gives)  $\Delta\lambda = \frac{1000 \times 10^3}{3 \times 10^8} \times 656.3$  (nm) = 2.19 (nm)  $\checkmark$   
(allow C.E. for value of v from (ii)  
 $\lambda_{\text{galaxy}} = 656.3 + 2.19 = 658.5$  nm  $\checkmark$  (4)

(b) for the furthest point of the Universe, 
$$d = \frac{c}{H} \checkmark$$
  
age of Universe  $= \frac{d}{c} = \frac{1}{H} \checkmark$   
[or use of  $v = Hd$  and  $t = \frac{d}{v} \checkmark$   
if all started from same point  
 $t = \text{age of Universe} = \frac{1}{H} \checkmark$ ]  
assumption: that *H* remains constant \checkmark

<u>(3)</u> (7)

4		
(a)	Hertzsprung -Russell diagram to show: absolute magnitude scale from +15 to −10 ✓ temperature scale from 50 000 to 2500 (K) ✓ main sequence drawn correctly ✓ giants and dwarfs shown in correct areas ✓	(4)
(b)	Alnitak : helium (absorption)Sirius : hydrogen Balmer (absorption) linesSun : metals (absorption)Antares : molecular bands	(2)
(c)	reference to $P = \sigma A T^4 \checkmark$ class M (Antares) cooler than class O (Alnitak) $\checkmark$ but same brightness, therefor cooler star bigger $\checkmark$ so Antares has larger surface area $\checkmark$	max <u>(3)</u> (9)
5 (a)(i)	supernova: star whose luminosity increase enormously due to it exploding ✓	
(a)(ii)	neutron star: star with the density of nuclear matter $\checkmark$	
(a)(iii)	black hole: an object whose escape velocity is greater than speed of light ✓	(3)
(b)	$\left(\text{use of } R = \frac{2GM}{c^2} \text{ gives}\right)  R = \frac{2 \times 6.67 \times 10^{-11} \times 10 \times 2 \times 10^{30}}{\left(3 \times 10^8\right)^2} \checkmark$	
	$= 2.96 \times 10^4 \text{ m} \checkmark$	<u>(2)</u> (5)

Quality of Written Communication (Q1(c)(i) and Q4(c))
$$\checkmark$$
(2)(2)(2)